

19. Costa, E. *et al.* *IAU Circ. No.* 6576 (1997).
 20. Yoshida, A. *et al.* *IAU Circ. No.* 6593 (1997).
 21. Groot, P. *et al.* *IAU Circ. No.* 6574 (1997).
 22. Landolt, A. UBVR photometric standard stars in the magnitude range 11.5–16.0 around the celestial equator. *Astron. J.* **104**, 340–376 (1992).
 23. Hakkila, J., Myers, J. M., Stidham, B. J. & Hartmann, D. H. A computerized model of large-scale visual interstellar extinction. *Astron. J.* (submitted).
 24. Johnson, H. L. Astronomical measurements in the infrared. *Annu. Rev. Astron. Astrophys.* **4**, 191–206 (1966).
 25. Luu, J. & Jewitt, D. Color diversity among the Centaurs and Kuiper Belt Objects. *Astron. J.* **112**, 2310–2318 (1996).
 26. Paczynski, B. *Variable Stars and the Astrophysical Returns of Microlensing Surveys* (ed. Ferlet, R. and Maillard, J.-P.) (Proc. 12th AIP Colloq., in the press).
 27. Trevese, D., Pittella, G., Kron, R. G., Koo, D. C. & Bershady, M. A survey for faint variable objects in SA 57. *Astron. J.* **98**, 108–116 (1989).
 28. Kochanski, G. P., Tyson, J. A. & Fischer, P. Flickering faint galaxies: few and far between. *Astron. J.* (in the press).
 29. Schechter, P. An analytic expression for the luminosity function for galaxies. *Astrophys. J.* **203**, 297–306 (1976).
 30. Pedersen, H. *et al.* Optical candidates for the 1978 November 19 gamma-ray burst source. *Astrophys. J.* **270**, L43–L47 (1983).
 31. Pedersen, H. & Hansen, J. A quasar in the 1978 November 19 gamma-ray burst error box. *Astrophys. J.* (submitted).
 32. Boer, M. *et al.* ROSAT detection and high precision localization of X-ray sources in the November 19, 1978 gamma-ray burst error box. *Astrophys. J.* (in the press).
 33. In't Zand, J. *et al.* *IAU Circ. No.* 6969 (1997).
 34. Greiner, J. & Heise, J. *IAU Circ. No.* 6570 (1997).
 35. Piro, L. *et al.* *IAU Circ. No.* 6570 (1997).
 36. Walsh, D. *et al.* Spectroscopy of 26 QSO candidates from the Jodrell Bank 966-MHz survey. *Mon. Not. R. Astron. Soc.* **211**, 105–109 (1984).
 37. Barthelmy, S. *et al.* Progress with the real-time GRB coordinates distribution Network (BACODINE). *AIP Proc.* **384**, 580–584 (1996).
 38. McNamara, B. *et al.* Ground-based gamma-ray burst follow-up efforts: results from the first two years of the BATSE/COMPTEL/NMSU Rapid Response Network. *Astrophys. J. Suppl. Ser.* **103**, 173–181 (1996).
 39. Vrba, F., Hartmann, D. & Jennings, M. Deep optical counterpart searches of gamma-ray burst localizations. *Astrophys. J.* **446**, 115–149 (1995).
 40. Vrba, F. Searches for gamma-ray burst counterparts: current status and future prospects. *AIP Proc.* **384**, 565–574 (1996).
 41. Luginbuhl, C., Vrba, F., Hudec, R., Hartmann, D. & Hurley, K. Results from the USNO quiescent optical counterpart search of IPN³ GRB and optical transient localizations. *AIP Proc.* **384**, 676–679 (1996).
 42. McNamara, B. *et al.* Ground-based γ -ray burst follow-up efforts: the first three years of the BATSE/COMPTEL/NMSU γ -Ray Burst Rapid Response Network. *AIP Proc.* **384**, 680–684 (1996).
 43. Castro-Tirado, A., Brandt, S., Lund, N. & Guziy, A. Optical follow-up of gamma-ray bursts observed by WATCH. *AIP Proc.* **307**, 404–407 (1994).
 44. Klose, S. Search for an optical counterpart of the source of GRB911001. *Astrophys. J.* **446**, 357–360 (1995).
 45. Schaefer, B. *et al.* Rapid searches for counterparts of GRB 930131. *Astrophys. J.* **422**, L71–L74 (1994).
 46. Park, H. S. *et al.* Limits on Real-Time Optical Emission from Gamma-Ray Bursts Measured by the GROCE Experiment. *Astrophys. J.* (submitted).
 47. Lee, B. *et al.* Results from GROCE: a real-time search for gamma-ray burst optical counterparts. *Astrophys. J.* (submitted).
 48. Castro-Tirado, A. *et al.* *IAU Circ. No.* 6598 (1997).
 49. Frail, D. *et al.* *IAU Circ. No.* 6576 (1997).
 50. Miller, H. R. & Noble, J. C. The microvariability of blazars and related AGN. *ASP Conf. Ser.* **110**, 17–29 (1996).

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Momentum creation by vortices in superfluid ³He as a model of primordial baryogenesis

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The Universe contains much more matter than antimatter, which is probably the result of processes in the early Universe in which baryon number was not conserved. These processes may have occurred during the electroweak phase transition, when elementary particles first acquired mass^{1–4}. It is impossible to study

directly processes relevant to the early Universe, because of the extreme energies involved. One is therefore forced to investigate laboratory systems with analogous phase transitions. Much of the behaviour of superfluid ³He is analogous to that predicted within the standard model of the electroweak interaction⁵. Superfluids and liquid crystals have already been used to investigate cosmic-string production^{6–11}; here we describe experiments on ³He that demonstrate the creation of excitation momentum (which we call momentogenesis) by quantized vortices in the superfluid. The underlying physics of this process is similar to that associated with the creation of baryons within cosmic strings, and our results provide quantitative support for this type of baryogenesis.

To explain the creation of matter in the early Universe we begin by recalling the relationship $E^2 = m_0^2 c^4 + p^2 c^2$ between energy E and momentum p for relativistic particles with rest mass m_0 (c is the velocity of light). Dirac realized that the square root of this equation, $E = \pm (m_0^2 c^4 + p^2 c^2)^{1/2}$ produced particles with both positive and negative energy. This led to his famous picture of the vacuum state (the state with no real particles) as one in which all the negative energy states are full and all the positive energy states are empty. A real particle is then created as shown in Fig. 1a by excitation of a particle from the ‘Dirac sea’ of negative-energy states to a positive-energy state. But such processes create matter and antimatter in equal amounts, as the appearance of a hole in the negative energy states is interpreted as the simultaneous creation of an antiparticle.

The net creation of matter in the form of baryons such as protons and neutrons requires processes in which antiparticles are not simultaneously created and thus baryon number is not conserved. In the ‘standard model’ the baryon number is classically conserved but can be violated by quantum-mechanical effects known generically as ‘chiral anomalies’. The process leading to matter creation is called ‘spectral flow’ and can be pictured as a process in which particles flow from negative energies to positive energies under the influence of an external force. In this way real observable positive-energy particles can appear without simultaneous creation of antiparticles. Figure 1b illustrates a simple example of spectral flow occurring for massless particles with electric charge q (in units of the charge of the electron, e) moving in a magnetic field. The application of an electric field E leads to the production of particles from the vacuum at the rate

$$\dot{n} = \partial_\mu j^\mu = \frac{q^2}{4\pi^2} \mathbf{E} \cdot \mathbf{B} \quad (1)$$

per unit volume, where j is the particle current fourvector; factors of ($e\hbar$) have been absorbed in the definition of the electric and magnetic fields. This is an anomaly equation for the production of particles from the vacuum of the type found by Adler¹² and by Bell and Jackiw¹³ in the context of neutral pion decay. We see that for particle creation it is necessary to have an asymmetric branch of the dispersion relation $E(p)$ which crosses the axis from negative to positive energy. We call such a branch a zero mode branch; a spectrum of this type was first found for vortex core excitations in a superconductor¹⁴.

Similar zero mode branches exist on a cosmic electroweak string (also known as a Z-string), which is a structure of the Higgs field that may have been produced during the electroweak phase transition. The Higgs field gives the particles mass outside the string core. This field vanishes on the string axis and the fermions (quarks and leptons) occurring in the core of the string behave like massless one-dimensional particles. Spectral flow on a Z-string leading to production of baryons from the vacuum with conservation of electric charge is illustrated in Fig. 1c. Motion of the string across a background electromagnetic field¹⁵ or the de-linking of two linked loops^{16,17} provides a mechanism for cosmological baryogenesis¹⁸ and could lead to the presence of antimatter in cosmic rays¹⁹.

We should point out that baryon-number violation is only one

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ingredient in a cosmological baryogenesis model. The other ingredients are that the system has to be out of thermal equilibrium, and charge and charge-parity conjugation symmetries should be violated; these conditions are known as the 'Sakharov conditions' for cosmological baryogenesis²⁰. In condensed matter the analogous symmetry breaking is provided by rotation (for ³He) or a magnetic field (for superconductors), and disequilibrium is provided by the motion of vortex lines. We shall see that excitation momentum is the analogue of baryon number.

The superfluidity of ³He is due to the formation of bound pairs of ³He atoms known as Cooper pairs. There are two main superfluid phases, A and B, which have different Cooper-pair wavefunctions. At the lowest temperature, the superfluid vacuum state is obtained in which all the atoms are Cooper-paired. By analogy with the Dirac picture of the vacuum (Fig. 1a) the Cooper-paired atoms are seen as filling all the negative energy states and are separated from the world of normal particles by the Cooper-pair binding energy. The Cooper-pair wavefunction (usually called the order parameter) which is responsible for this energy gap is thus analogous to the Higgs field in electroweak theory, as it is the finite particle mass produced by the Higgs field that results in the energy gap in electroweak theory between the negative- and positive-energy states in Fig. 1a. At higher temperatures in ³He some atoms are excited from the ground state by the breaking of Cooper pairs and these normal (non-superfluid) excitations interact with the order parameter in a way that is similar to the interaction of the standard model particles with the electroweak gauge and Higgs fields. Thus the behaviour of leptons and quarks on fixed gauge and Higgs field backgrounds can be modelled by the behaviour of ³He excitations on fixed order parameter backgrounds such as quantized vortices. In ³He a zero-mode branch exists for particles in the core of quantized vortices (Fig. 2a). A physically important charge in ³He-A, ³He-B and superconductors, which (like baryonic charge in the standard model) is not conserved due to the anomaly, is excitation momentum. Spectral flow along the zero-mode branch leads to an additional 'lift' force on a moving vortex (Fig. 2b, the analogue of Fig. 1c).

This is most easily seen for the doubly quantized continuous vortex (winding number $N = 2$) in the A-phase of superfluid ³He, which is the closest analogue of a Z-string. The Cooper pairs in ³He-

A have angular momentum \hbar and locally all the pairs have this angular momentum aligned along a direction \hat{l} . The $N = 2$ vortex is characterized by a continuous distribution of the order parameter vector \hat{l} as shown in Fig. 2c. The time- and space-dependent \hat{l} vector associated with the motion of the vortex produces a force on the excitations equivalent to that of an 'electric field' $\mathbf{E} = k_F \partial_t \hat{l}$ and a 'magnetic field' $\mathbf{B} = k_F \nabla \times \hat{l}$ acting on particles of unit charge, where $k_F = p_F/\hbar$ and p_F is the Fermi momentum. Equation (1) can then be applied to calculate the rate at which left-handed quasi-particles and right-handed quasiholes are created by spectral flow. As both types of excitation have momentum $p_F \hat{l}$, excitation momentum is created at a rate:

$$\partial_t \mathbf{P} = \frac{1}{2\pi^2} \int d^3 r p_F \hat{l} (\mathbf{E} \cdot \mathbf{B}) \quad (2)$$

However, total linear momentum must be conserved. Therefore equation (2) means that, in the presence of a time-dependent texture, momentum is transferred from the superfluid ground state (analogue of vacuum) to the heat bath of excitations forming the normal component (analogue of matter).

Integration of the anomalous momentum transfer in equation (2) over the cross-section of the moving vortex gives the loss of linear momentum and thus the additional force per unit length acting on the vortex due to spectral flow:

$$\mathbf{F}_{sf} = \partial_t \mathbf{P} = \pi \hbar N C_0 \hat{z} \times (\mathbf{v}_n - \mathbf{v}_l) \quad (3)$$

Here \hat{z} is the direction of the vortex, $C_0 = k_F^3/3\pi^2$, \mathbf{v}_l is the velocity of the vortex line, \mathbf{v}_n is the heat bath velocity, and N is the winding number of the vortex.

The insets to Fig. 3a show how the force on a vortex is measured experimentally. A uniform array of vortices is produced by rotating the whole cryostat, and oscillatory superflow perpendicular to the rotation axis is produced by a vibrating diaphragm, while the normal fluid (thermal excitations) is clamped by viscosity. The velocity \mathbf{v}_l of the vortex array is determined by the overall balance of forces acting on the vortices, conventionally written as^{21,22}

$$n_s \pi \hbar N [\hat{z} \times (\mathbf{v}_l - \mathbf{v}_s) + d_\perp \hat{z} \times (\mathbf{v}_n - \mathbf{v}_l) + d_\parallel (\mathbf{v}_n - \mathbf{v}_l)] = 0 \quad (4)$$

where $n_s(T)$ and \mathbf{v}_s are the density and velocity of the superfluid

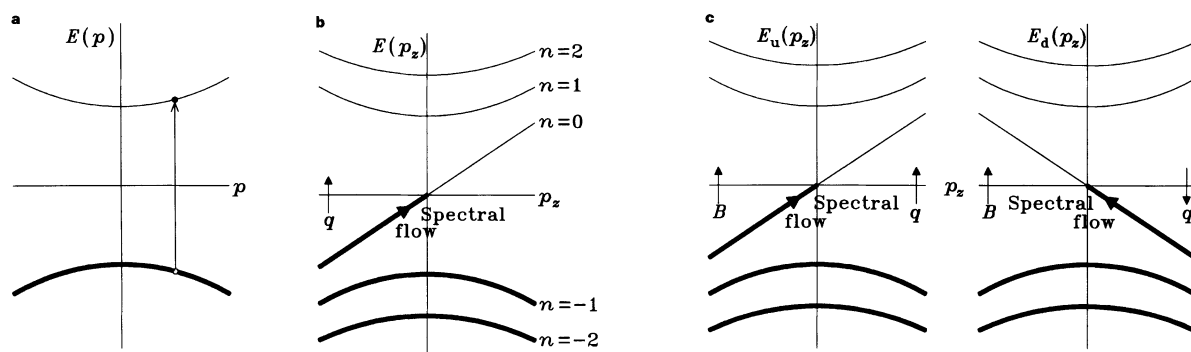


Figure 1 **a**, Particles and antiparticles of energy E and momentum p in the Dirac picture; the thick line shows occupied negative-energy states. Promotion of a particle from negative energy to positive energy creates a particle-antiparticle pair from the vacuum. **b**, Spectrum of massless right-handed particles in a magnetic field \mathbf{B} along z ; the thick lines show the occupied negative-energy states. The right-handed chirality of the particle means that their spin is aligned with their linear momentum. Motion of the particles in the plane perpendicular to \mathbf{B} is quantized into the Landau levels shown. The free motion is thus effectively reduced to one-dimensional motion along \mathbf{B} with momentum p_z . Because of the chirality of the particles the lowest ($n = 0$) Landau level, for which $E = cp_z$, is asymmetric: it crosses zero only in one direction. If we now apply an electric field \mathbf{E} along z , particles of charge q are pushed from negative- to positive-energy levels according to the equation of motion $\dot{p}_z = qE$, and the whole Dirac sea (see

text) moves up, creating particles and electric charge from the vacuum. This motion of the particles along the 'anomalous' branch of the spectrum is called spectral flow. The rate of particle production is proportional to the density of states at the Landau level, which is $\propto q^2 |\mathbf{B}|$, so that the rate of production of particles from the vacuum $\propto q^2 \mathbf{E} \cdot \mathbf{B}$. **c**, Anomalous branches of the spectrum of u-quarks ($q = +2/3$; left panel) and d-quarks ($q = -1/3$; right panel) in the core of a Z-string. There are anomalous branches for electrons ($q = -1$) and neutrinos ($q = 0$) also. The neutrino and u-quark propagate in one direction along the string while the electron and d-quark propagate in the opposite direction. For every electron produced, two u-quarks and one d-quark are created so that there is no net production of electric charge but the baryon number B increases by one ($2 \text{ u-quarks} + 1 \text{ d-quark} = 1 \text{ proton} = 1 \text{ baryon}$).

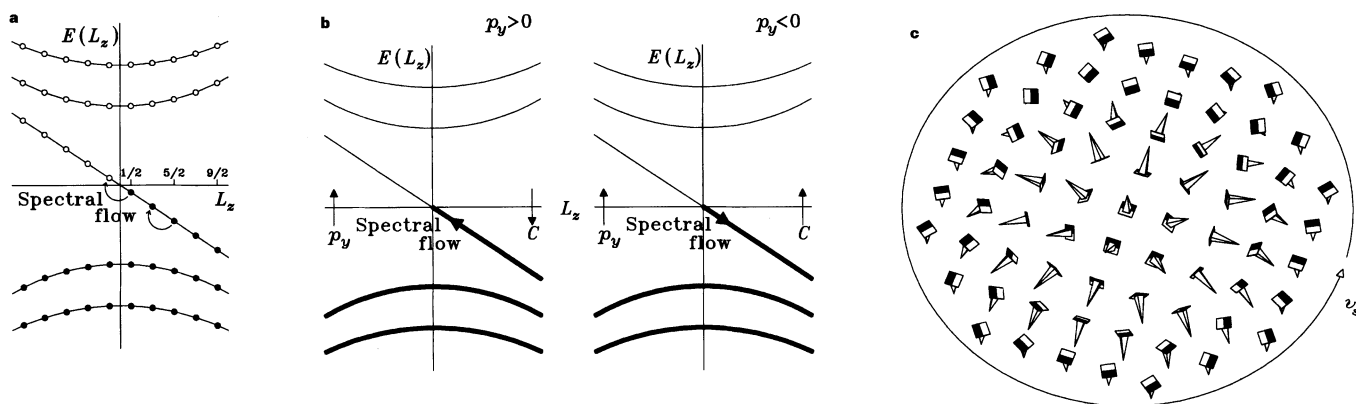


Figure 2 a, Spectrum of particles in the core of vortices in ^3He and superconductors; filled circles show occupied states. The only difference from the Z-string is that the anomalous branch, $E_0(p_z, L_z) = -L_z \hbar \omega_0(p_z)$, of the spectrum crosses zero as a function of the discrete angular momentum $\hbar L_z$, where L_z is half an odd integer. Typically the interlevel distance $\hbar \omega_0$ is very small compared to the characteristic energy scales in superconductors and Fermi superfluids, and can be comparable to the level width \hbar/τ resulting from the scattering of core excitations by free excitations in the heat bath outside the core. If $\omega_0 \tau < 1$, the levels overlap and spectral flow is allowed. This type of spectral flow is analogous to 'hopping conduction' in a solid; it is assisted rather than impeded by collisions. **b**, Such spectral flow is induced by the motion of the vortex with respect to the heat bath. If the vortex moves with velocity v_x along x , the rate of change of excitation angular momentum in the moving core is $L_z = x p_y = v_x p_y$. The levels

cross zero at a rate $-L_z/\hbar$ leading to spectral flow in opposite senses for $p_y > 0$ and $p_y < 0$ as indicated. As the left-handed quasiparticles created for $p_y > 0$ are equal in number to the right-handed quasiparticles created for $p_y < 0$ there is no net production of chiral charge C , but excitation momentum is created at a rate $p_y = -v_x p_y^2/\hbar$, independent of the sign of p_y (shown for negative v_x). **c**, An example of an $N = 2$ continuous vortex in $^3\text{He-A}$. The cones indicate the local direction of the order parameter vector $\hat{\mathbf{l}}$, which is parallel to the angular momentum of the Cooper pairs, and is analogous to the direction of the isotopic spin in electroweak theory. Rotation of the cones about $\hat{\mathbf{l}}$ indicates a change in phase of the order parameter; note the 4π phase change around the perimeter of the diagram, corresponding to two quanta of anticlockwise circulation. Four topologically different types of $^3\text{He-A}$ vortices have been observed²⁵.

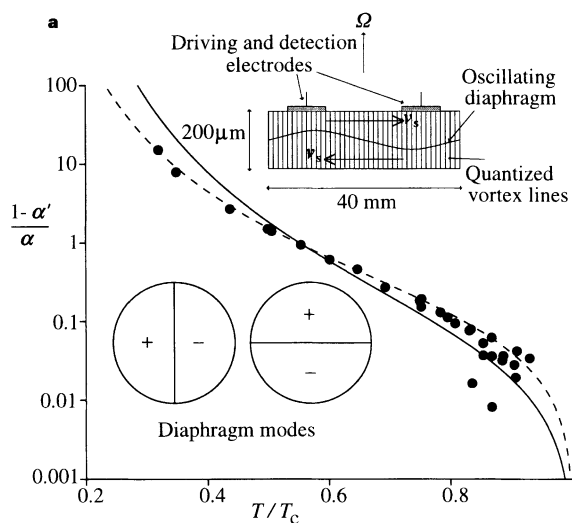
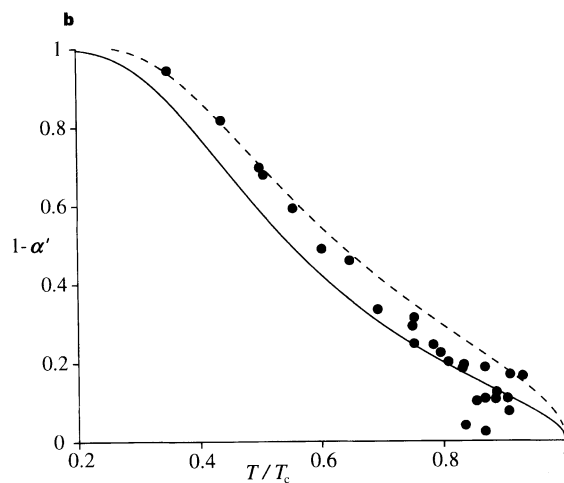


Figure 3 a, The experimental cell is shown in the top inset. The aluminized Kapton film diaphragm separates two disk-shaped regions of superfluid ^3He , each $100\mu\text{m}$ thick. The roof of the cell has six electrodes set into it by means of which the oscillations of the diaphragm may be driven and detected electrostatically. In the oscillating modes of interest the motion of the diaphragm displaces the superfluid as indicated, while the normal component of the fluid (the heat bath) is clamped by its high viscosity. Rotation at angular velocity Ω about a vertical axis produces vortices normal to the diaphragm. These vortices produce additional dissipation proportional to $\alpha\Omega$ and coupling between two orthogonal modes of the diaphragm, with displacement patterns shown in the bottom inset, proportional to $(1 - \alpha')\Omega$. The measured spectral flow parameters are related to the parameters in ref. 22 by $\alpha = B n_n/2n$ and $\alpha' = B' n_n/2n$.



Main figure, the measured $(1 - \alpha')/\alpha$ for $^3\text{He-B}$ at 20 bar is compared with the theoretical value $(1 - \alpha')/\alpha \approx \omega_0 \tau$. The temperature dependence of $\omega_0 \tau$ is not very well known, as it is sensitive to the details of the quasiparticle scattering. It can be estimated by the formula²¹ $\omega_0 \tau = (\partial \Delta(T)/\partial k_B T) \exp(\Delta(T)/k_B T)$. The solid line is for $a = 0.067$. The fit is much improved if an effective energy gap 0.6 of the bulk value is assumed; the dashed line is for $\Delta(T) = 0.6 \Delta_{\text{bulk}}(T)$ and $a = 0.214$. But the justification for such an assumption is not clear. **b**, The experimental parameter $(1 - \alpha')$ for $^3\text{He-B}$ at 20 bar is compared with the 'spectral flow' prediction of eq. (5): $(1 - \alpha') \approx n_s(T)/(n \tanh(\Delta(T)/2k_B T))$ (solid line). Note that this fit is independent of the relaxation parameter $\omega_0 \tau$. The dashed line is for $\Delta(T) = 0.6 \Delta_{\text{bulk}}(T)$, and fits much better below $0.6T_c$. This fit shows that spectral flow is less strongly suppressed at low temperatures than the full energy gap would suggest.

component. The first term is the Magnus force, which appears when the vortex moves with respect to the superfluid, and the terms with $(\mathbf{v}_n - \mathbf{v}_l)$ represent the nondissipative transverse and frictional longitudinal forces, proportional to dimensionless parameters d_\perp and d_\parallel respectively, which appear if the vortex moves with respect to the normal fluid. The diaphragm has to provide a force equal and opposite to the Magnus force to drive the superfluid. Measurement of the damping of the diaphragm resonance and of the coupling between the two orthogonal modes illustrated in Fig. 3a enables both d_\perp and d_\parallel to be deduced.

For the A-phase, the spectral flow force in equation (3) combines with the Iordanskii force²¹ to give $d_\perp \approx (C_0 - n_n)/n_s$, where $n_n(T) = n - n_s(T)$ is the density of the normal component, with n the total density. As $C_0 \approx n$, one has $d_\perp \approx 1$. Our ³He-A experiments made at 29.3 bar and $T > 0.82T_c$, where T_c is the superfluid transition temperature, are consistent with this within experimental uncertainty: we find that $|1 - d_\perp| < 0.005$ (ref. 23).

Any Fermi superfluid becomes similar to the A-phase in the vicinity of the vortex core. However, in ³He-B the spacing $\hbar\omega_0$ between bound states on the anomalous branch in Fig. 2a becomes larger than the broadening $\hbar\tau^{-1}$ due to the lifetime τ at low T (ref. 21). There should thus be a transition from full spectral flow as $T \rightarrow T_c$ to totally suppressed spectral flow as $T \rightarrow 0$, a further aspect of the theory which can be tested experimentally. The interpolation formula as a function of the relaxation parameter $\omega_0\tau$ and the spectral flow parameter $C_0 \approx n$ (refs 21, 24) can be written in the form

$$d_\parallel - i(1 - d_\perp) = \frac{1}{\alpha + i(1 - \alpha')} = \frac{n}{n_s} \frac{\omega_0\tau}{1 + i\omega_0\tau} \tanh \frac{\Delta(T)}{2k_B T} \quad (5)$$

where $i = (-1)^{\frac{1}{2}}$ and the α parameters are what are directly measured experimentally²². The experimental results are compared with equation (5) in Fig. 3. The agreement is excellent in view of the approximations in the theory.

Our results thus show that the chiral anomaly is relevant for the interaction of condensed matter vortices (analogous to strings) with fermionic excitations (analogous to quarks and leptons); this gives a firmer footing to chiral anomaly calculations for baryogenesis on non-trivial backgrounds of the Higgs field such as cosmic strings. Superfluid ³He is the most complex field-theoretical system available in the laboratory, and is therefore the closest experimental analogue of the field-theoretical foundations of the physics of the early Universe. It is characteristic of nonlinear theories that experiment often reveals qualitatively new phenomena that would be hard to find by consideration of the equations alone. We may therefore hope that imaginative experiments on superfluid ³He will generate new ideas in cosmology. □

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1. Dolgov, A. D. Nongut baryogenesis. *Phys. Rep.* **222**, 309–386 (1992).
2. Turok, N. in *Formation and Interaction of Topological Defects* (eds Davis, A. C. & Brandenberger, R.) 283–301 (Plenum, New York, 1995).
3. Vilenkin, A. & Shellard, E. P. S. *Cosmic Strings and Other Topological Defects* (Cambridge Univ. Press, 1994).
4. Hindmarsh, M. B. & Kibble, T. W. B. Cosmic Strings. *Rep. Prog. Phys.* **58**, 477–562 (1995).
5. Volovik, G. E. & Vachaspati, T. Aspects of ³He and the standard electroweak model. *Int. J. Mod. Phys. B* **10**, 471–521 (1996).
6. Zurek, W. H. Cosmological experiments in superfluid helium. *Nature* **317**, 505–508 (1985).
7. Chuang, I., Durrer, R., Turok, N. & Yurke, B. Cosmology in the laboratory: defect dynamics in liquid crystals. *Science* **251**, 1336–1342 (1991).
8. Bowick, M. J., Chander, L., Schiff, E. A. & Srivastava, A. M. The cosmological Kibble mechanism in the laboratory: string formation in liquid crystals. *Science* **263**, 943–945 (1994).
9. Hendry, P. C., Lawson, N. S., Lee, R. A. M., McClintock, P. V. E. & Williams, C. D. H. Generation of defects in superfluid ³He as an analogue of the formation of cosmic strings. *Nature* **368**, 315–317 (1994).
10. Ruutu, V. M. H. et al. Vortex formation in neutron irradiated ³He as an analogue of cosmological defect formation. *Nature* **382**, 334–336 (1996).
11. Bäuerle, C., Bunkov, Yu. M., Fisher, S. N., Godfrin, H. & Pickett, G. R. Laboratory simulation of cosmic string formation in the early Universe using superfluid ³He. *Nature* **382**, 332–334 (1996).
12. Adler, S. Axial-vector vertex in spinor electrodynamics. *Phys. Rev.* **177**, 2426–2438 (1969).
13. Bell, J. S. & Jackiw, R. A PCAC puzzle: $\pi^+ \rightarrow \gamma\gamma$ in the σ -model. *Nuovo Cim. Ser.* **1060A**, 47–61 (1969).
14. Caroli, C., de Gennes, P. G. & Matricon, J. Bound fermion states on a vortex line in a type II superconductor. *Phys. Lett.* **9**, 307–309 (1964).
15. Witten, E. Superconducting strings. *Nucl. Phys. B* **249**, 557–592 (1985).

16. Vachaspati, T., Field, G. B. Electroweak string configurations with baryon number. *Phys. Rev. Lett.* **73**, 373–376 (1994); **74**, 1258(E) (1995).
17. Garriga, J. & Vachaspati, T. Zero modes on linked strings. *Nucl. Phys. B* **438**, 161–181 (1995).
18. Barriola, M. Electroweak strings produce baryons. *Phys. Rev. D* **51**, 300–304 (1995).
19. Starkman, G. D. & Vachaspati, T. Galactic cosmic strings as sources of primary antiprotons. *Phys. Rev. D* **53**, 6711–6714 (1996).
20. Sakharov, A. Violation of CP invariance, C asymmetry, and baryon asymmetry of the Universe. *Pis'ma Zh. Eksp. Teor. Fiz.* **5**, 32–35 (1967); *JETP Lett.* **5**, 24–27 (1967).
21. Kopnin, N. B., Volovik, G. E. & Parts, Ü. Spectral flow in vortex dynamics of ³He-B and superconductors. *Europhys. Lett.* **32**, 651–656 (1995).
22. Bevan, T. D. C. et al. Vortex mutual friction in rotating superfluid ³He-B. *Phys. Rev. Lett.* **74**, 750–753 (1995); **74**, 3092(E) (1995).
23. Manninen, A. J. et al. Vortex mutual friction, orbital inertia and history dependent textures in rotating superfluid ³He-A. *Phys. Rev. Lett.* **77**, 5086–5089 (1996).
24. Stone, M. Spectral flow, Magnus force and mutual friction via the geometric optics limit of Andreev reflection. *Phys. Rev. B* **54**, 13222–13229 (1996).
25. Parts, Ü et al. Phase diagram of vortices in superfluid ³He-A. *Phys. Rev. Lett.* **75**, 3320–3323 (1995).

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Morphogenesis of shapes and surface patterns in mesoporous silica

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The synthesis of inorganic materials with complex form, using surfactant assemblies as supramolecular templates^{1–13}, has ramifications in areas as diverse as large-molecule catalysis, the formation of semiconductor nanostructures, biomolecule separations, the development of medical implants and the morphogenesis of skeletal forms¹⁴. Here we describe a procedure for the synthesis of hexagonal mesoporous silica that produces a remarkable array of shapes, surface patterns and channel plans. Our reaction conditions favour curved morphologies including toroidal, disk-like, spiral and spheroidal shapes. We use scanning electron microscopy to catalogue the basic topologies and surface patterns, and transmission electron microscopy to establish the relationship between morphology and the underlying mesostructure. Polarized optical microscopy enables us to identify a connection between optical anisotropy in these structures and the periodic porous mesostructure. We propose that the morphogenesis of these shapes and surface patterns can be rationalized in terms of the growth of a silicate liquid-crystal embryo^{15,16} with a hexagonal cross-section that, under different initial reaction conditions, is subject to increasing degrees of curvature.

We synthesized these mesoporous silica bodies under quiescent aqueous acidic conditions using cetyltrimethylammonium chloride (CTACl) as surfactant template and tetraethylorthosilicate (TEOS) as silica precursor^{10,13,17}. The materials obtained are ordinary in that they have the well documented hexagonal symmetry channel structure of mesoporous MCM-41², as signalled by the commonality of their diagnostic powder X-ray diffraction (PXRD) pattern^{13,17}. On the other hand, scanning electron microscopy (SEM) images reveal that the morphologies are extraordinary in terms of their diverse and remarkable shapes and impressive range of curvatures (Fig. 1). High-resolution SEM images of the exterior surfaces of the curved mesoporous silica morphologies show that they tend to fall into two classes: some are smooth at the 200–500-Å length scale but with signs of stepped and spiral growth, whereas others display faceted corrugated patterns that emanate radially or